



ليزر عملي

الكورس الاول / التجربة الثالثة

المرحلة الثالثة

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# Spectroscopy

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## **Purpose**

Finding the wavelength of the laser beam

## **Devices used**

Source (neon)

diffraction grating

Spectroscopy

## **Theory**

The diffraction grating is used to study and analyze the neon spectrum to see that it consists of several lines (several wavelengths) only one of them has the largest profit in the laser device, where these lengths can be measured and the wavelength used in the work of the laser device is found by measuring the angle of each line and at the specified level.

The spectroscope consists of three main parts: the socket, which is a non-nucleated lens, whose function is to form a parallel beam of light, a window from the lens and the endoscope, and it is a regular astronomical telescope whose function is to receive the spectrum and see it through the eyepiece and the base scale, with which we can read angles using a vernier

We notice from the following figure the most important transitions between the energy levels of the neon atom, the first

table gives the type of transition and the wavelength of the red color associated with it, and the second table gives the five most important transitions in laser work because of the presence of a good profit value, and we notice that the largest profit occurs in the 632.8 nm wavelength, which is the wavelength For helium-neon lasers

611.8 nm	10
629.8 nm	20
632.8 nm	100
635.2 nm	6
640.1 nm	34

Transition	Wavelength(nm)
3s <sup>2</sup> →2p <sup>1</sup>	730.5
3s <sup>2</sup> →2p <sup>2</sup>	640.1
3s <sup>2</sup> →2p <sup>3</sup>	635.2
3s <sup>2</sup> →2p <sup>4</sup>	632.8
3s <sup>2</sup> →2p <sup>5</sup>	629.8
3s <sup>2</sup> →2p <sup>6</sup>	611.8
3s <sup>2</sup> →2p <sup>7</sup>	604.6
3s <sup>2</sup> →2p <sup>8</sup>	593.9

### Accounts and discussion

1. Set up the spectrometer and turn on the red neon lamp so that the slit of the spectrometer appears clearly lit and matches the wire

2. The notch is fixed in the place designated for it and the telescope rotates until the first line of the spectrum is seen from the left and the angle is read, then the telescope rotates to the right through the direct main line until the first line of the spectrum from the right and the angle is read, then the second line from the left and right and so on.
3. The diffraction angle is calculated  $\theta_n$
4. Calculates the wavelength

$$\lambda = \frac{\sin \theta_n}{N}$$